

Big Dithers are Crucial for LSST Survey Uniformity

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Initial Results for Task CWG-2: "Full-sky Simulations with the Operations Simulator"

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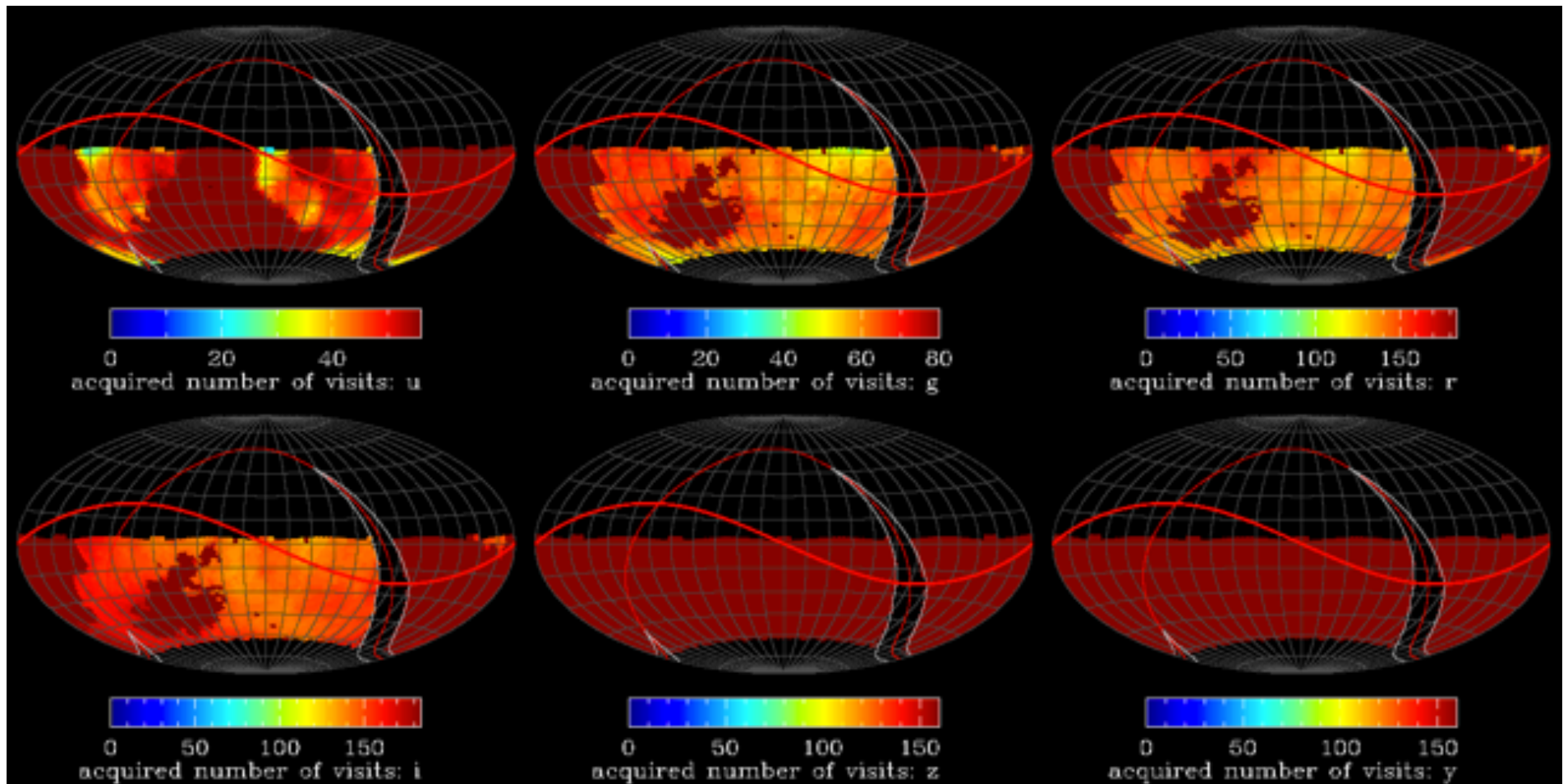
The LSST Operations Simulator

Non-trivial survey design task to cover
20,000 square degrees divided into
2000 pointings, each observed
~150 times in each of 6 filters.

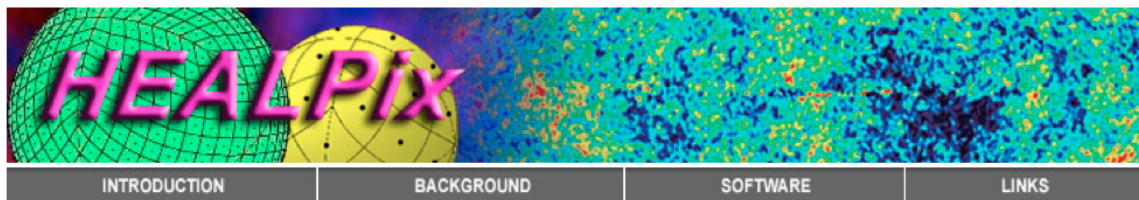
Feasibility shown via Operations Simulator (OpSim) that
decides where the telescope should point next and
simulates realistic atmospheric and telescope behavior
meta-data over the full 10-year survey (over 1 million
exposures).

Challenge: *OpSim does not resolve LSST pointings, which
are circles covering a hexagonal tiling of the sphere.*

Task CWG-2: "Full-sky Simulations with the Operations Simulator"



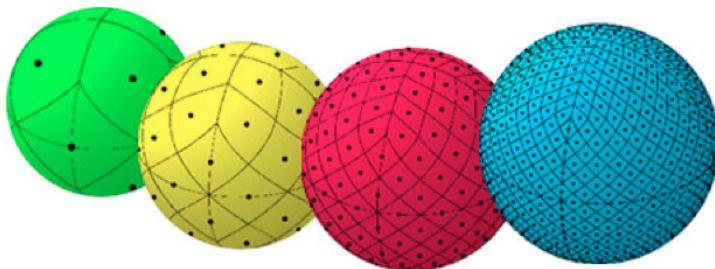
OpSim output shows number of visits ($\approx \text{depth}^2$) for LSST pointings centered on hexagonal tiling of the sky



Introduction

HEALPix is an acronym for **H**ierarchical **E**qual **A**rea iso**L**atitude **P**ixelization of a sphere. As suggested in the name, this pixelization produces a subdivision of a spherical surface in which each pixel covers the same surface area as every other pixel. The figure below shows the partitioning of a sphere at progressively higher resolutions, from left to right. The green sphere represents the lowest resolution possible with the HEALPix base partitioning of the sphere surface into 12 equal sized pixels. The yellow sphere has a HEALPix grid of 48 pixels, the red sphere has 192 pixels, and the blue sphere has a grid of 768 pixels (~ 7.3 degree resolution).

Another property of the HEALPix grid is that the pixel centers, represented by the black dots, occur on a discrete number of rings of constant latitude, the number of constant-latitude rings is dependent on the resolution of the HEALPix grid. For the green, yellow, red, and blue spheres shown, there are 3, 7, 15, and 31 constant-latitude rings, respectively.

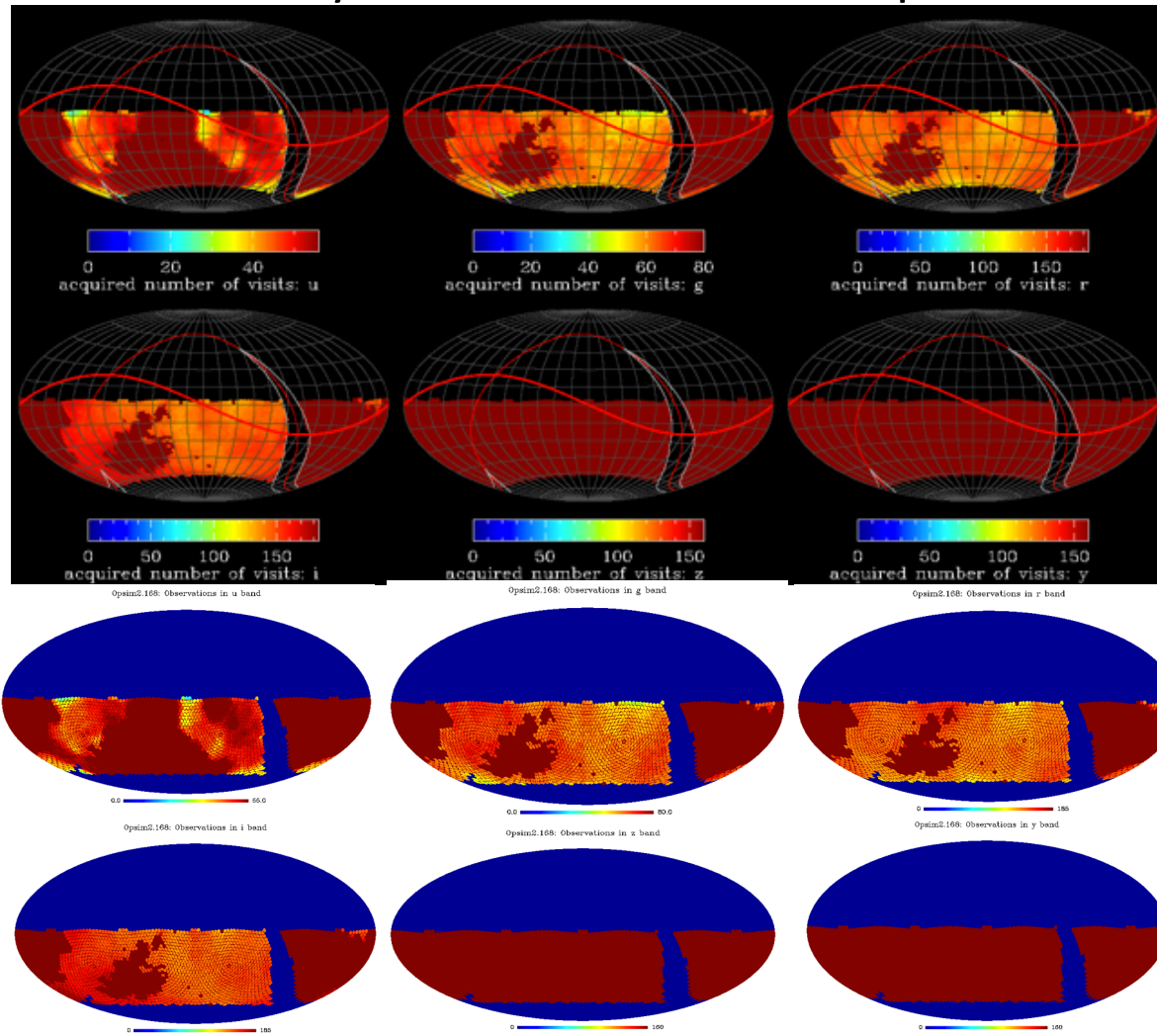


HEALPix turns the sphere into a vector of equal-area pixels ordered by declination. Good for fast spherical harmonic transforms.

We use HEALPix at level 128 with 196608 pixels on the sphere to resolve each \sim circular LSST FOV into ~ 50 pixels.

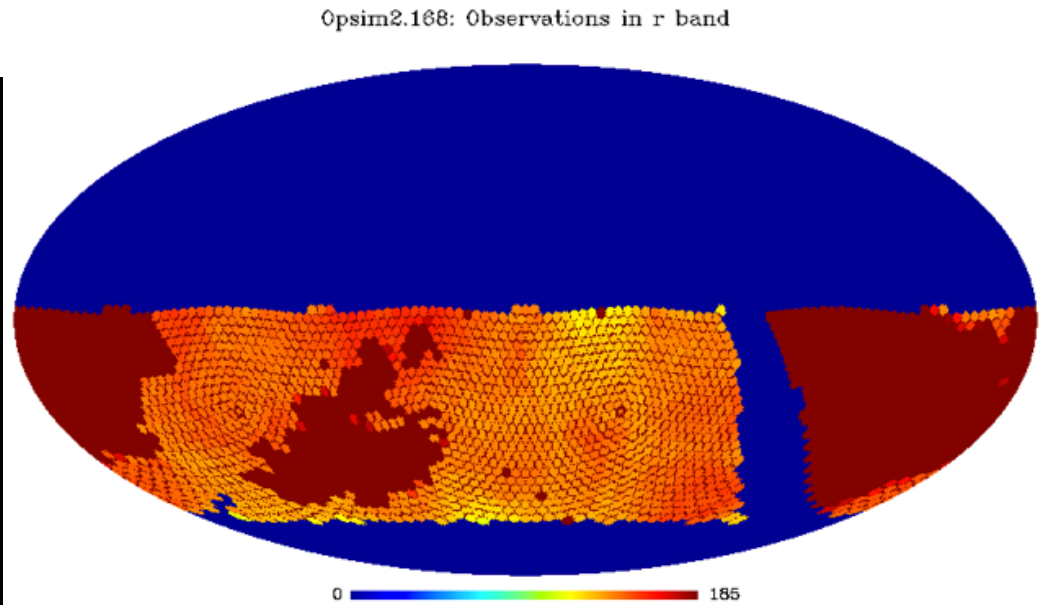
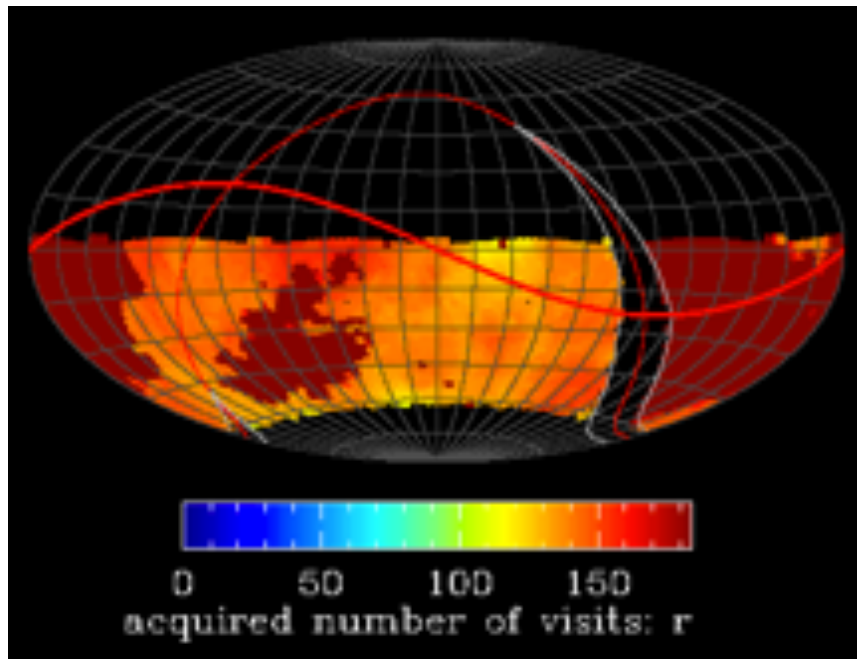
see <http://healpix.jpl.nasa.gov/index.shtml>, ArXiv:astro-ph/0409513, ArXiv:astro-ph/9905275

Task CWG-2: "Full-sky Simulations with the Operations Simulator"



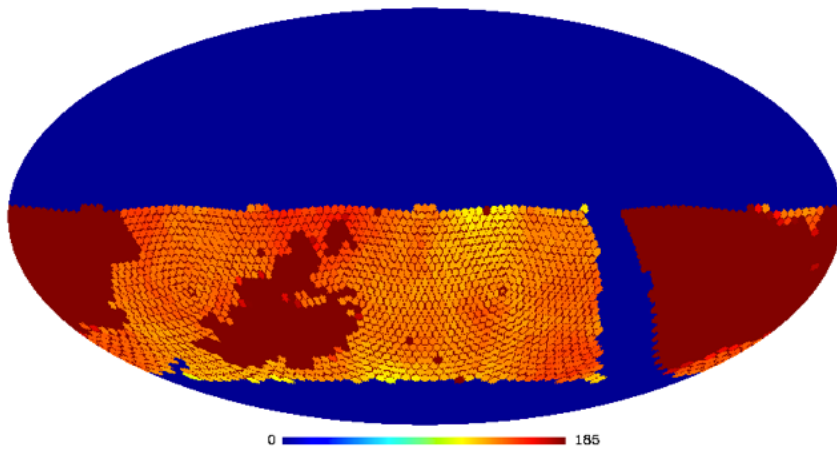
We match this, except for shift from Aitoff to Mollweide projection

Task CWG-2: "Full-sky Simulations with the Operations Simulator"

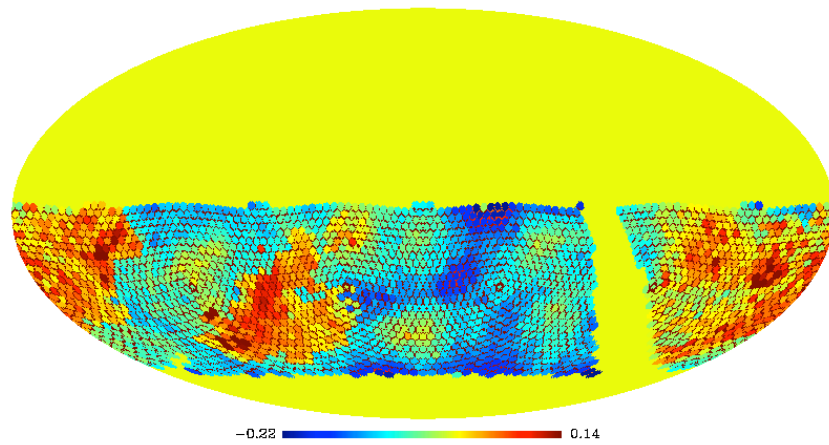


We match this in detail, but resolving LSST FOV with HEALPIX reveals expected honeycomb pattern of overlapping regions (16% of survey area) that receive double coverage

Opsim2.168: Observations in r band



Delta N / N_ave - 27.500 Cutoff

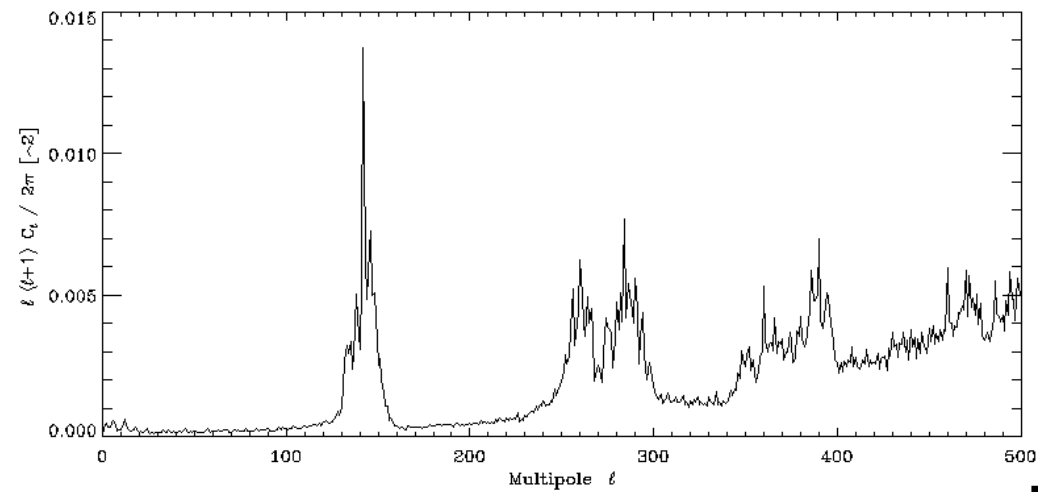
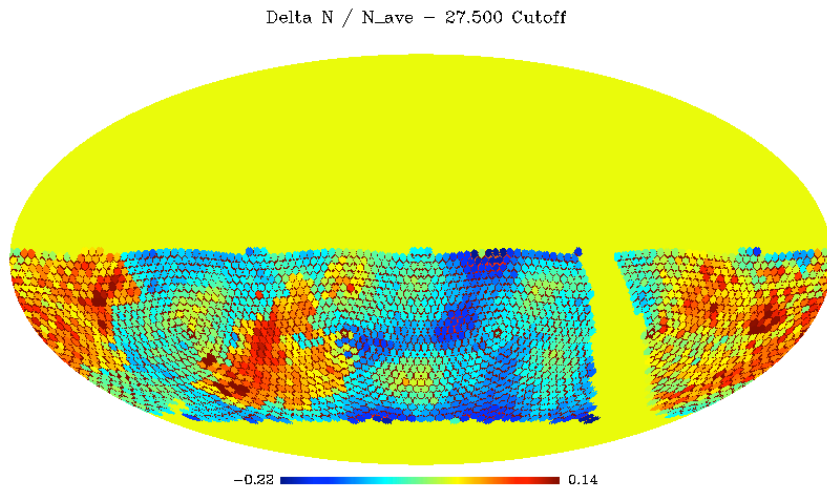


But we don't care about r-band depth!

Convert it to predicted galaxy $\delta N / \langle N \rangle$ i.e. fake large-scale structure (right skymap).

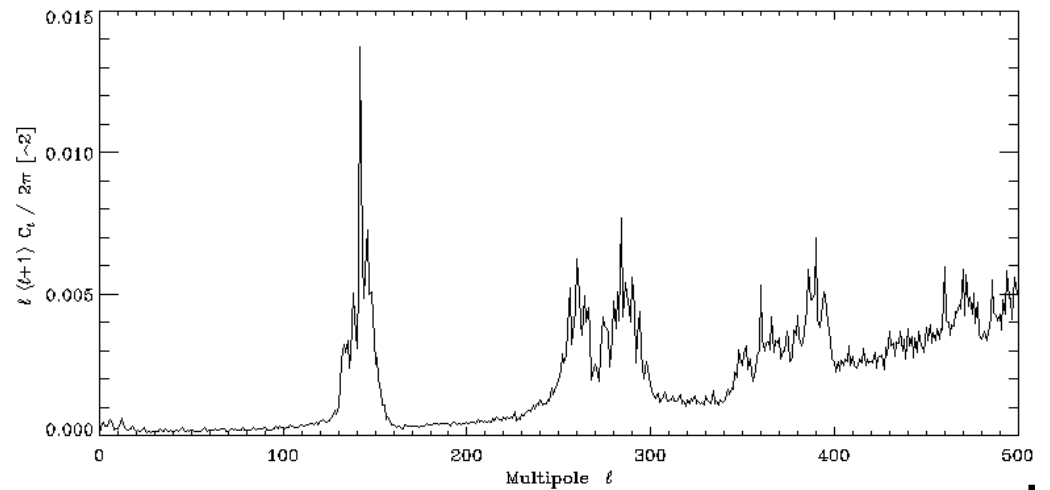
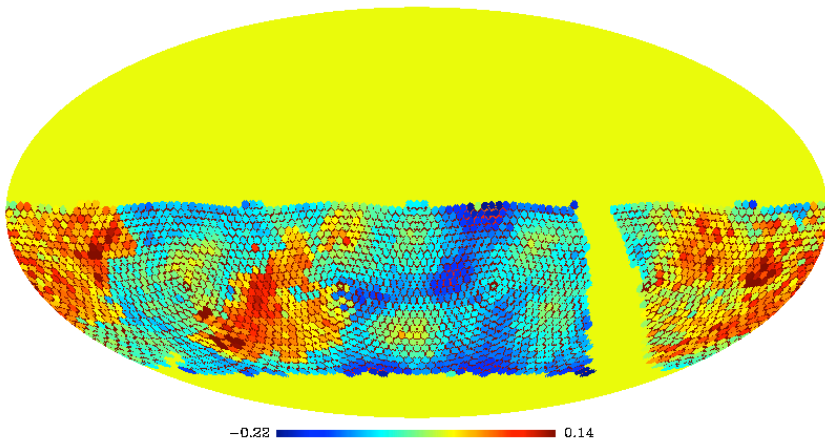
We currently use:

- MUSYC R-band number counts
- Official OpSim 5 sigma point source depths with optimal coadds (much smarter than 2" mag limits but surely imperfect)
- Fleming (1995) formula to model incompleteness & Malmquist bias

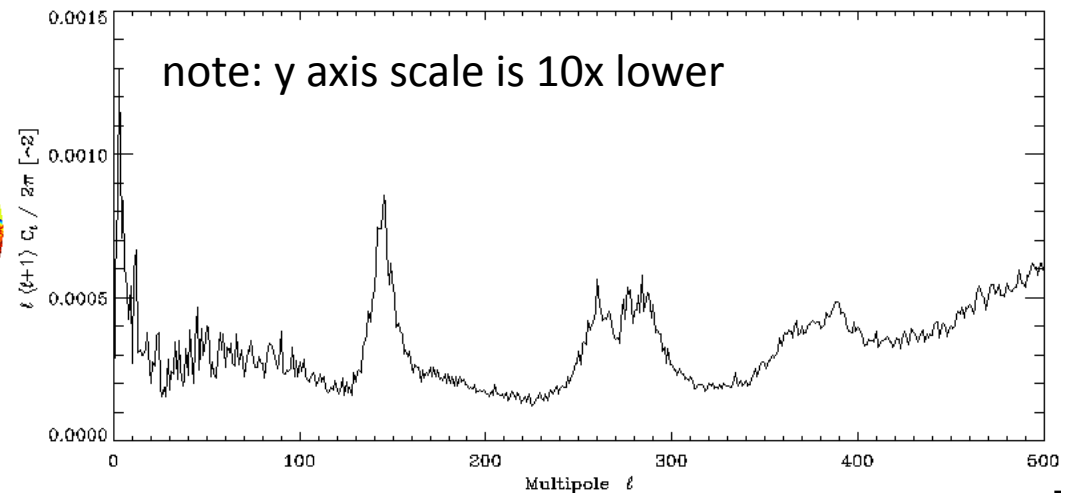
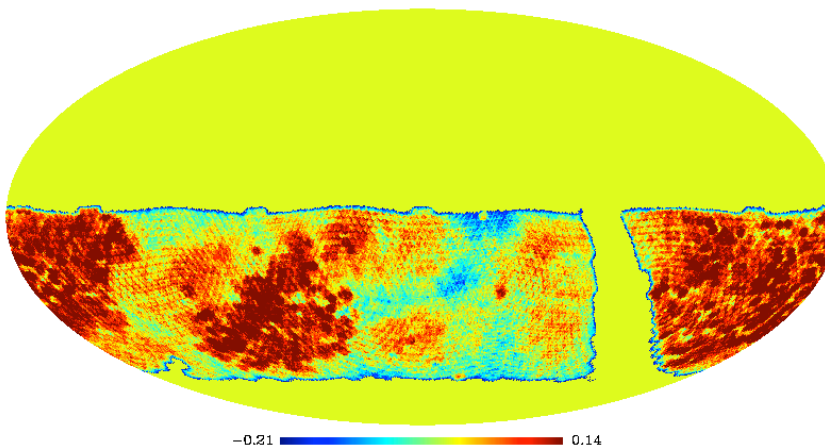


HEALPIX shows the angular power spectrum of $\delta N / \langle N \rangle$. Features at $100 < \ell < 300$ cause systematic errors in B.A.O. studies of dark energy

Delta N / N_ave - 27.500 Cutoff



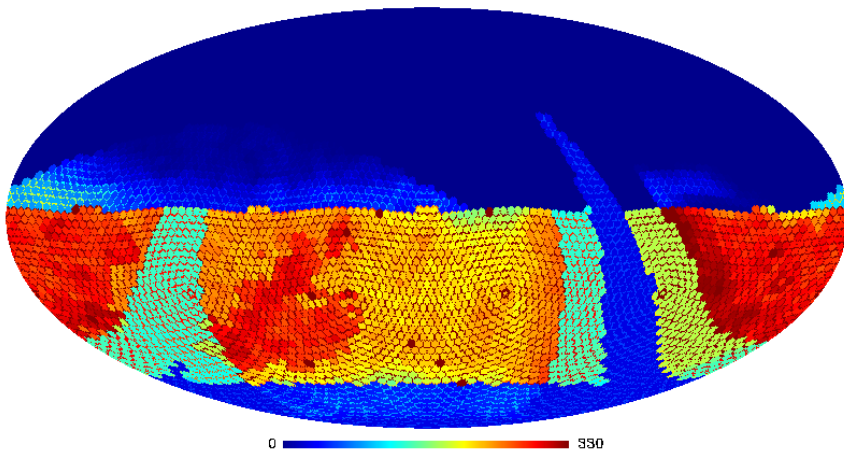
Delta N / N_ave - 27.500 Cutoff



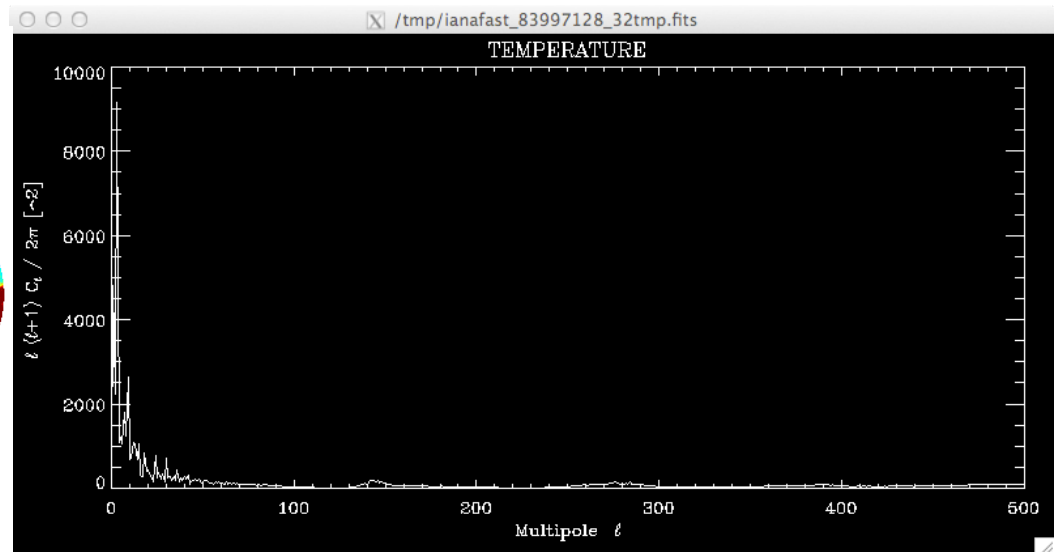
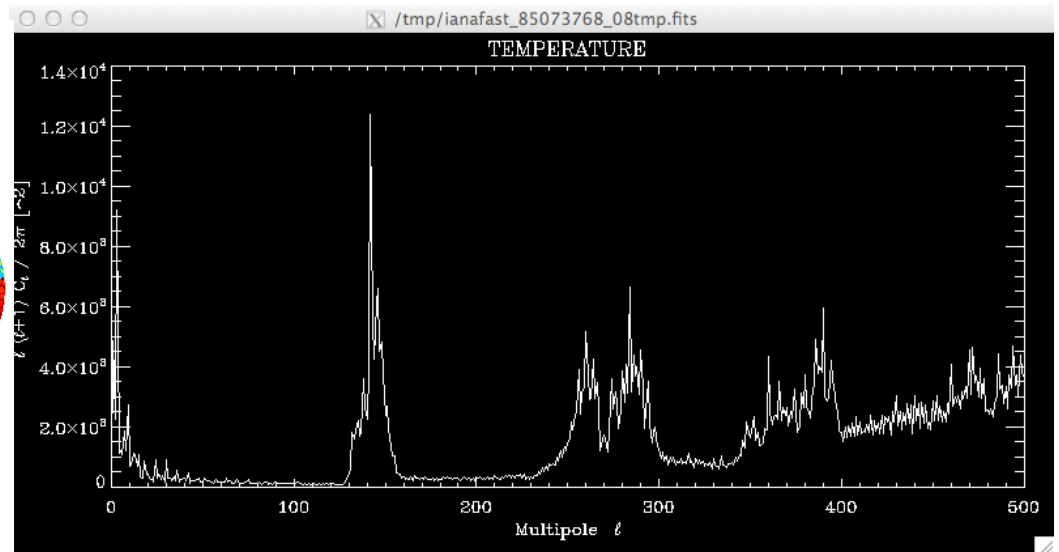
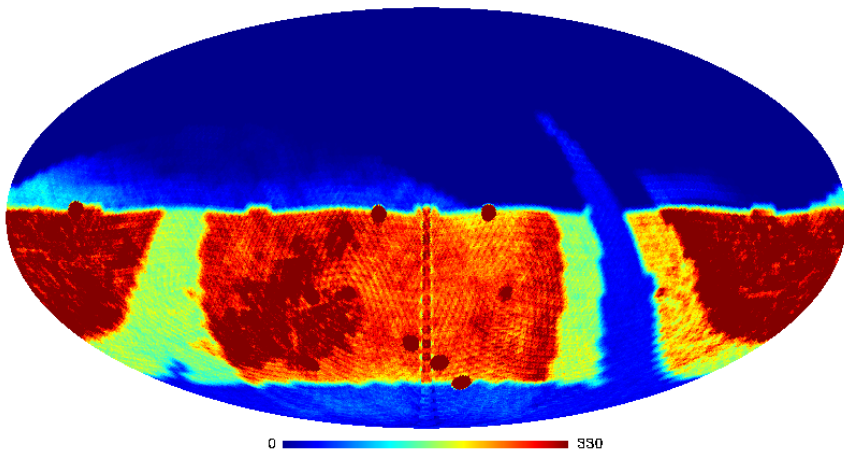
HEALPIX shows the angular power spectrum of $\delta N / \langle N \rangle$. Features at $100 < l < 300$ cause systematic errors in B.A.O. studies of dark energy, **but large dithers reduce these by a factor of 10!!!**

Task CWG-2: "Full-sky Simulations with the Operations Simulator"

Opsim2.168: Observations in r band



Opsim2.168: Observations in r band (Dithering)



Now including mini-surveys & better color table, but C_ℓ are just for "Main Survey".
Large (~FOV) dithers produce far more uniform coverage but are not desirable for Deep Drilling Fields.

Status Report on Task CWG-2: "Full-sky Simulations with the Operations Simulator"

✓ Use HEALPIX (at level $n_{\text{side}}=128$) to subdivide each LSST pointing into about 50 sub-pointings, allowing us to track statistics such as number and depth of exposures in overlapping regions between pointings. Implement a large-dithering scheme as a post-processing of OpSim.

• Study the cadence of exposure of the overlapping regions to determine if the $\sim 2\times$ improvement causes a significant increase in the number of Type Ia supernovae with sufficient cadence for cosmological studies. As of now, only Deep Drilling Field supernovae are considered acceptable for cosmology. Underway, with David Cinabro & Rahul Biswas.

✓ Use the built-in spherical harmonic transform capabilities of HEALPIX to study the angular power spectrum of depth variations and various sources of systematics using the simulated observational metadata (atmospheric conditions etc.) provided by OpSim. This will reveal the level of artificial power on the BAO scale.

• Work (ongoing) with Jones, Yoachim and the OpSim team to test alterations to the LSST survey to explore a variety of dithering schemes including rotational dithers to determine their effects on LSS systematics at the BAO scale and larger scales.

Bottom line: Big dithers are undesirable for the Deep Drilling Fields but crucial for Main Survey uniformity.